

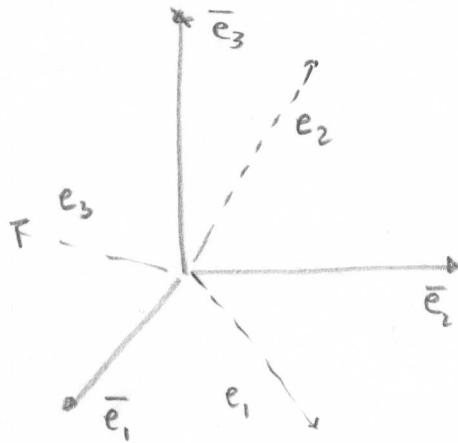
Rigid Body Motion

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Rigid body: distance between points is fixed

6 degrees of freedom: 3 translation + 3 rotation.

Consider a body fixed at a point  $\rightarrow$  rotation



$(\bar{e}_1, \bar{e}_2, \bar{e}_3) \rightarrow$  fixed space frame

$(e_i(+), e_j(+), e_k(+)) \rightarrow$  moving body frame

Both coordinate systems are orthogonal:  $\bar{e}_i \cdot \bar{e}_j = \delta_{ij}$

$$e_i \cdot e_j = \delta_{ij}$$

$\Rightarrow$  Find an orthogonal matrix  $R = (R_{ik})$  such that

$$(*) \quad e_i = R_{ik}(+) \bar{e}_k = \sum_{k=1}^3 R_{ik}(+) \bar{e}_k$$

$$\begin{aligned} \text{Proof: } e_i \cdot e_j &= \delta_{ij} = (R_{ik} \bar{e}_k) \cdot (R_{jm} \bar{e}_m) = R_{ik} R_{jm} \underbrace{\bar{e}_k \cdot \bar{e}_m}_{\delta_{km}} \\ &= R_{ik} R_{jk} = (R^T R)_{ij} \Rightarrow R \text{ is orthogonal} \end{aligned}$$

$$\text{From } (*): R_{ik} = e_i \cdot \bar{e}_k$$

We are now ready to introduce angular velocity:

[ We can express any point  $\vec{r}$  in space frame or body frame:  $\vec{r} = \vec{r}_i(t) \vec{e}_i = r_k e_k(t)$  and  $\vec{r}_k(t) = r_i R_{ik}(t)$

$$\frac{d\vec{r}}{dt} = \frac{d\vec{r}_i}{dt} \vec{e}_i = r_i \frac{de_i}{dt} = r_i \frac{dR_{ik}}{dt} \vec{e}_k ]$$

For the  $(e_1, e_2, e_3)$ , the body frame, in particular

$$\begin{aligned} \frac{de_i}{dt} &= \frac{dR_{ik}}{dt} \vec{e}_k = \frac{dR_{ik}}{dt} \vec{R}_{kj}^{-1} e_j \quad \text{as } \vec{e}_k = \vec{R}_{kj}^{-1} e_j \\ &=: \tilde{\omega}_{ij} e_j \end{aligned}$$

with  $\tilde{\omega}_{ij} = \dot{R}_{ik} (\vec{R}^{-1})_{kj} = \dot{R}_{ik} R_{jk}$  (as  $R^T = \vec{R}^{-1}$ )

It is easy to see that  $\omega$  is antisymmetric:  $\tilde{\omega}_{ik} = -\tilde{\omega}_{ki}$

Proof:  $R_{ik} R_{jk} = \delta_{ij} \Rightarrow \dot{R}_{ik} R_{jk} + R_{ik} \dot{R}_{jk} = 0$   
 $\Rightarrow \tilde{\omega}_{ij} + \tilde{\omega}_{ji} = 0 \quad \square$

# Analytical Dynamics

Lecture 7

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$$\tilde{\omega} = \begin{pmatrix} 0 & \omega_3 - \omega_2 \\ -\omega_3 & 0 & -\omega_1 \\ \omega_2 - \omega_1 & 0 \end{pmatrix}$$

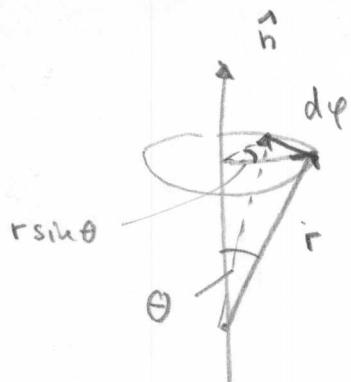
and we can write with  $\omega = \begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix}$   
meaning  $\omega = \omega_i e_i$

⇒ 
$$\frac{d e_i}{dt} = \tilde{\omega} \times e_i$$

Ex: 
$$\frac{d e_1}{dt} = \tilde{\omega}_{1j} e_j = \tilde{\omega}_{12} e_2 + \tilde{\omega}_{13} e_3 = \omega_3 e_2 - \omega_2 e_3$$

$$\omega \times e_1 = (\omega_1 e_1 + \omega_2 e_2 + \omega_3 e_3) \times e_1 = -\omega_2 e_3 + \omega_3 e_2$$

## Interpretation



Consider small rotation  $d\varphi$  around axis  $\hat{n}$

$$|dr| = |r||d\varphi| \sin \theta$$

$$\Rightarrow dr = d\varphi \times r \quad d\varphi = \hat{n} |d\varphi|$$

$$\Rightarrow \frac{dr}{dt} = \omega \times r \quad \text{with } \omega = \frac{d\varphi}{dt}$$

⇒  $\omega$  is the instantaneous angular velocity.

We can now express e.g.  $T$  the kinetic energy using  $\omega$ :

$$T = \frac{1}{2} \sum_i m_i \dot{r}_i^2 = \frac{1}{2} \sum_i m_i (\omega \times r_i) \cdot (\omega \times r_i)$$

$$= \frac{1}{2} \sum_i m_i ((\omega \cdot \omega)(r_i \cdot r_i) - (r_i \cdot \omega)^2)$$

$$(*) = \frac{1}{2} \omega_i I_{ik} \omega_k \quad \text{where } \underline{\text{the inertia tensor }} I$$

is defined as

$$I_{ik} = \sum_j m_j ((r_j \cdot r_j) \delta_{ik} - (r_j)_i (r_j)_k)$$

Let the rigid body have a density  $\rho(r)$ , then

$$I = \int d^3r \rho(r) \begin{pmatrix} y^2 + z^2 & -xy & -xz \\ -xy & x^2 + z^2 & -yz \\ -xz & -yz & x^2 + y^2 \end{pmatrix}$$

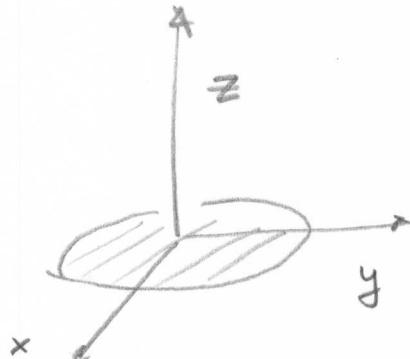
Note:  $I$  is symmetric, hence it can be diagonalized:

Find an orthogonal matrix  $M$  such that  $I' = M I M^T$  is diagonal. The axes which are aligned with the eigenvectors of  $I$  are called principal axes

Example: disc: (no mass in z-direction)

$$\Sigma = \frac{M}{\pi R^2}$$

$$I_{11} = \int_{\text{Disc}} y^2 \frac{M}{\pi R^2} dx dy$$

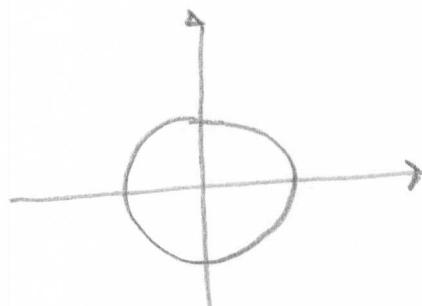


To compute this integral:

$$I_{22} = \int_{\text{Disc}} x^2 \frac{M}{\pi R^2} dx dy$$

$$I_{33} = I_{11} + I_{22} = \int_{\text{Disc}} x^2 + y^2 \frac{M}{\pi R^2} dx dy = \frac{M}{\pi R^2} \int_0^R r^3 dr \cdot 2\pi$$

$$= 2\pi \frac{M}{\pi R^2} \cdot \frac{R^4}{4} = \frac{MR^2}{4}$$



$$\Rightarrow I_{11} = I_{22} = \frac{MR^2}{4}$$

For the angular momentum:

$$L = \sum_j m_j (\mathbf{r}_j \times \mathbf{v}_j) = \sum_j m_j (\mathbf{r}_j \times (\omega \times \mathbf{r}_j))$$

$$= \sum_j m_j [\omega r_j^2 - (\mathbf{r}_j \cdot \omega) \mathbf{r}_j]$$

$$\text{e.g. } L_x = \sum_j m_j [\omega_x r_j^2 - (\mathbf{r}_j \cdot \omega) x_j]$$

$$= \sum_j m_j [\omega_x (x_j^2 + y_j^2 + z_j^2) - (x_j \omega_x + y_j \omega_y + z_j \omega_z) x_j]$$

$$= I_{xx} \omega_x + I_{xy} \omega_y + I_{xz} \omega_z$$

$$I_{xx} = \sum_j m_j (y_j^2 + z_j^2) \quad I_{xy} = \sum_j m_j (-x_j y_j)$$

$$I_{xz} = \sum_j m_j (-x_j z_j)$$

or, in general

$$L = I \omega$$

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## Analytical Dynamics

Most general motion:

overall translation superposed  
with a rotation

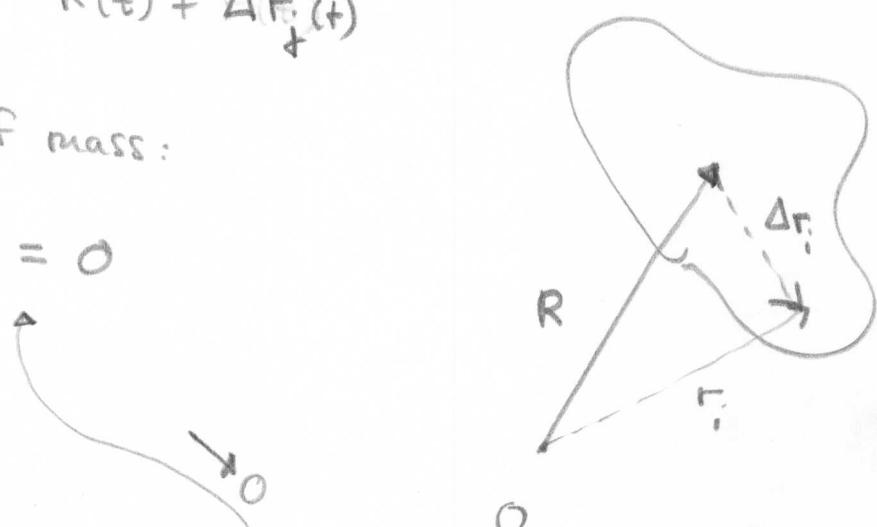
↳ useful: consider rotation around center of mass

$$\vec{r}_j(t) = \vec{R}(t) + \Delta\vec{r}_j(t)$$

For center of mass:

$$\sum_j m_j \Delta r_j = 0$$

⇒



$$T = \frac{1}{2} \sum_j m_j \dot{r}_j^2$$

$$= \sum_j m_j \left( \frac{1}{2} \dot{R}^2 + \dot{R} \cdot (\omega \times \Delta r_j) + \frac{1}{2} (\omega \times \Delta r_j)^2 \right)$$

$$= \frac{1}{2} M \dot{R}^2 + \frac{1}{2} \omega_i I_{ij} \omega_j$$

(Center of mass:  $\vec{R} = \frac{1}{M} \sum m_j \vec{r}_j$ )

$$\Rightarrow \sum_j m_j \Delta r_i = \sum_j m_j (\vec{r}_j - \vec{R}) = \sum_j m_j \vec{r}_j - \sum_j R m_j = 0$$

## Free motion of rigid body

$$\frac{dL}{dt} = 0 \quad \text{and} \quad L = L_k e_k^{\text{body axes}}$$

$$0 = \frac{dL_k}{dt} e_k + L_k \frac{de_k}{dt} = \frac{dL_k}{dt} e_k + L_k (\omega \times e_k) \quad (*)$$

Let body axes coincide with principal axes:

$$L_k = I_k \omega_k, \text{ so e.g. } L_1 = I_1 \omega_1$$

$$\omega = \omega_j e_j$$

Project (\*) on  $e_i$

$$-\frac{dL_i}{dt} = L_k (\omega_j e_j \times e_k) \cdot e_i \quad \text{if } i=1, \text{ we have}$$

$$-\frac{dL_i}{dt} = I_k \omega_k \omega_j (e_j \times e_k) \cdot e_i \quad \begin{aligned} e_2 \times e_3 &= e_1 \\ e_3 \times e_2 &= -e_1 \end{aligned}$$

$$0 = I_1 \dot{\omega}_1 + I_3 \omega_3 \omega_2 - I_2 \omega_2 \omega_3$$

$$0 = I_1 \dot{\omega}_1 + \omega_2 \omega_3 (I_3 - I_2)$$

⇒ Euler's Equations

$$I_1 \dot{\omega}_1 + \omega_2 \omega_3 (I_3 - I_2) = 0$$

$$I_2 \dot{\omega}_2 + \omega_3 \omega_1 (I_1 - I_3) = 0$$

$$I_3 \dot{\omega}_3 + \omega_1 \omega_2 (I_2 - I_1) = 0$$

Example: Symmetric top

$$I_1 = I_2 \neq I_3$$

$$I_1 \dot{\omega}_1 = \omega_2 \omega_3 (I_1 - I_3)$$

$$I_2 \dot{\omega}_2 = -\omega_1 \omega_3 (I_1 - I_3)$$

$$I_3 \dot{\omega}_3 = 0$$

⇒ Spin about the symmetric axis with constant  $\omega_3$

$$\Rightarrow \dot{\omega}_1 = \Omega \omega_2, \quad \dot{\omega}_2 = -\Omega \omega_1, \quad \Omega = \frac{\omega_3 (I_1 - I_3)}{I_1}$$

$$\Rightarrow (\omega_1, \omega_2) = \omega_0 (\sin \Omega t, \cos \Omega t)$$

(Precession about the  $e_3$  axis with frequency  $\Omega$ )

("wobble" of earth tricky to explain - more complicated)

Example: Stability of asymmetric top:

Assume  $I_1 \neq I_2 \neq I_3 \neq I_1$ . Spin completely about one of the principal axes:

$$\omega_1 = \Omega, \quad \omega_2 = \omega_3 = 0$$

What happens to a small perturbation:

$$\omega_1 = \Omega + \zeta_1, \quad \omega_2 = \zeta_2, \quad \omega_3 = \zeta_3$$

$$\Rightarrow I_1 \dot{\zeta}_1 = 0$$

$$I_2 \dot{\zeta}_2 = \Omega \zeta_3 (I_3 - I_1)$$

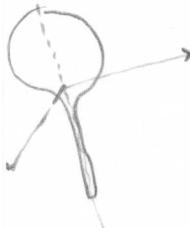
$$I_3 \dot{\zeta}_3 = \Omega \zeta_2 (I_1 - I_2)$$

$$\Rightarrow I_2 \ddot{\zeta}_2 = \frac{\Omega^2}{I_3} (I_3 - I_1) (I_1 - I_2) \zeta_2$$

$\underbrace{\hspace{10em}}$

If  $\alpha > 0 \Rightarrow$  unstable,  $\alpha < 0 \Rightarrow$  stable

Unstable:  $I_2 < I_1 < I_3$  or  $I_3 < I_1 < I_2$



Stable about axis with smallest and largest moment of inertia.